

# Standard Cosmological models

Cosmology: building representations of the  
“world”

~ one of the oldest activity of human kind.

→ prejudices can be strong (and hidden)

# Physical Cosmology...

Simple principles

+

Physical laws

- → a consistent picture
- → predictions
- → check/validation by observations

**... is a mature science.**

≠ **having basic principles for the universe**

→ infers physical laws

- “Perfect Cosmological Principle”

→ violates energy conservation

- “Tired light theory”

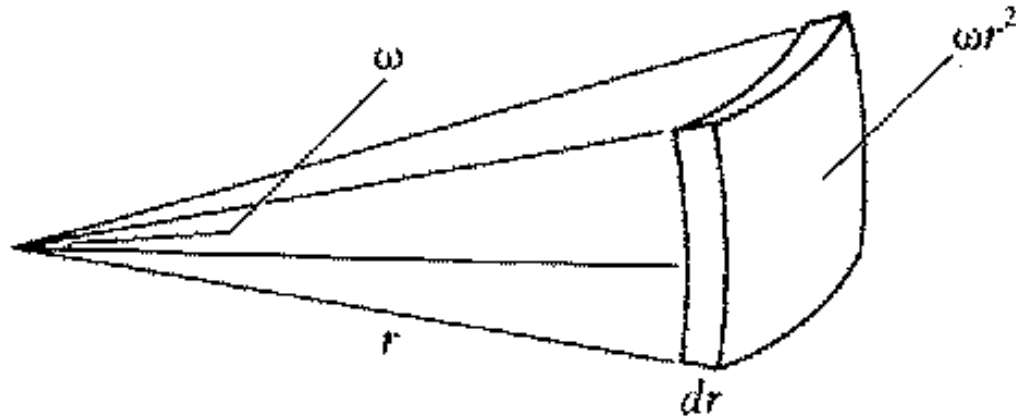
→  $\nu$  shift is due to some unknown physical process.

## However:

- introduction of the cosmological constant by Einstein (?)  
→ to have solutions under the starting static hypothesis.
- early universe physics  
→ advocates unknown physics (inflation, ...)

# Olbers paradox

Volume element for counts:



# Olbers paradox

Homogeneous medium of stars:

$$l = \frac{L}{4\pi r^2}$$

Number of stars between  $r$  and  $r + dr$ :

$$dN = n_* \times \omega \times r^2 \times dr$$

Number of stars between  $l$  and  $l + dl$ :

$$dN = -\frac{1}{2} n_* \omega \left( \frac{L}{4\pi} \right)^{3/2} \frac{dl}{l^{5/2}}$$

# Number counts:

$$N(> l) = \frac{1}{3} n_* \omega \left( \frac{L}{4\pi} \right)^{3/2} \frac{1}{l^{3/2}}$$

in term of magnitude ( $m = -2.5 \log(l) + \text{cste}$ )

$$\log(N(< m)) \propto 0.6m + \text{cste}$$

# Olbers paradox

Integrated luminosity:

$$\phi = \int_0^{+\infty} \frac{dN}{dl} l dl$$

$l^{-3/2}$  makes the integral diverging!

Something wrong among:

- The universe is homogeneous
- Universe is static and eternal
- Geometry of space is Euclidian geometry



In retrospect, now that we have reasonably convincing evidence that the universe really is expanding, it is easy to find reasons why a static universe is problematic.

J.Peebles

# Olbers paradox

## Energetics considerations:

- Finite volume  $\Rightarrow$
- Finite number of stars  $\Rightarrow$
- Finite amount of energy is available

# Olbers paradox

Timescale for energy exhaust:

Order of magnitude for the Sun:

$$L_{\odot} \sim 4 \cdot 10^{33} \text{ erg/s}$$

$$M_{\odot} \sim 2 \cdot 10^{33} \text{ g}$$

Efficiency of nuclear reactions  $\epsilon \sim 0.007$

$$\begin{aligned} \tau &\sim \frac{\epsilon M_{\odot} c^2}{L_{\odot}} = \frac{E_t}{\frac{dE_t}{dt}} \\ &\sim 3 \cdot 10^{18} \text{ s} \sim 10^{11} \text{ yr} \end{aligned}$$

=> the universe cannot remain identical for ever!

# Homogeneity

“The universe is homogeneous on large scale”

Einstein cosmological principle

Can and should be tested from observations.

$$\lim_{R \rightarrow \infty} \bar{\rho}(R) = cste$$

(necessary but not sufficient...)

From galaxies:

$$\log(N(m)) \propto 0.6m ?$$

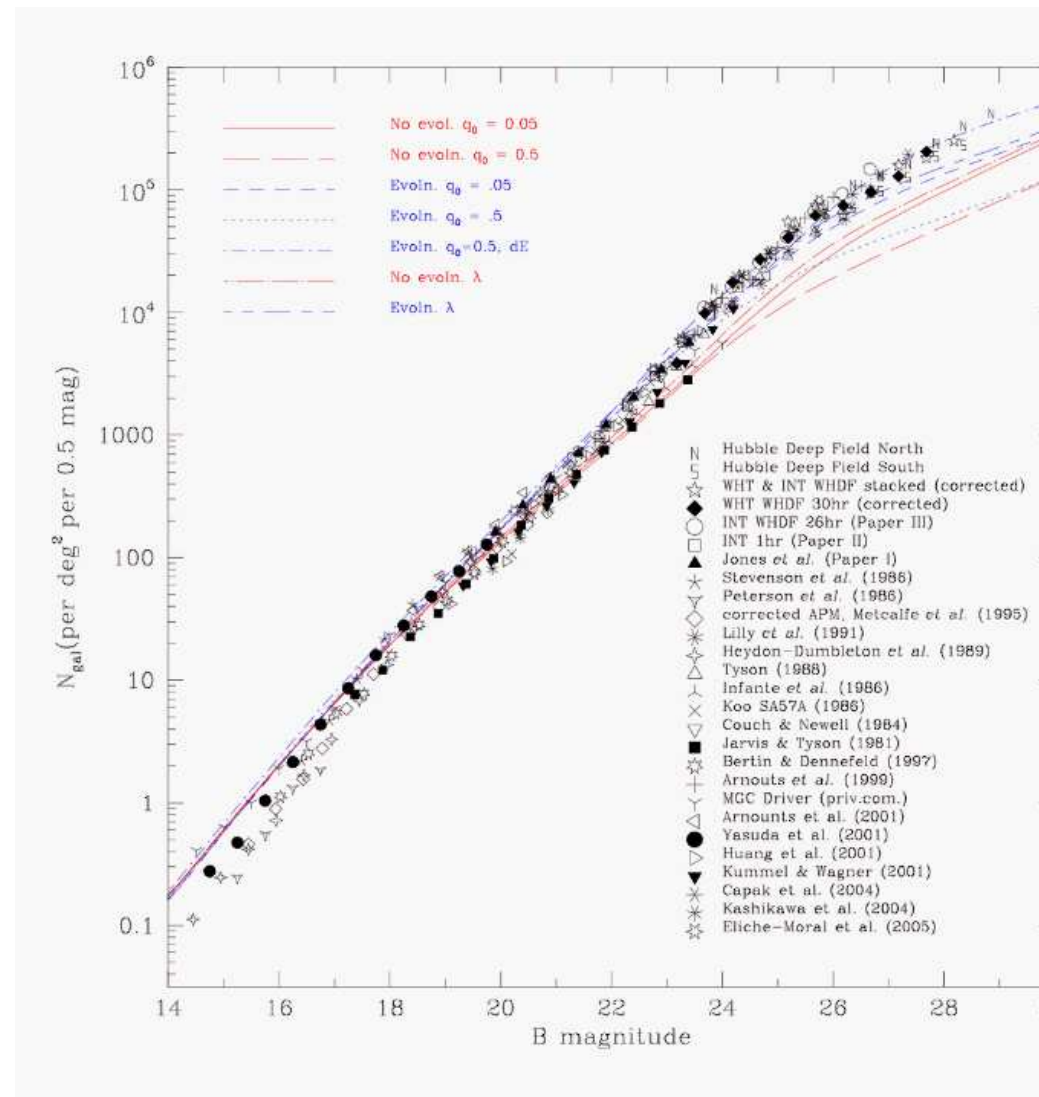
good indication.

Isotropy

+Copernic principle => homogeneity

# Homogeneity

From galaxies number counts:



# Geometry

- locally we can assign four coordinates to an event  
 $(x, y, z, t)$
- this does not prejudge of the global *shape* of the universe:  
Plane ? Sphere ? Torrus ?

3D Spherical universe: start from 4D  $(x, y, z, u)$

$$x^2 + y^2 + z^2 + u^2 = R^2$$

using spherical coordinates  $r^2 = x^2 + y^2 + z^2$   
from  $dl^2 = dx^2 + dy^2 + dz^2 + du^2$  and  $u^2 = R^2 - r^2$  one gets:

$$dl^2 = r^2(d\theta^2 + \sin^2 \theta d\phi^2) + \frac{dr^2}{1 - \left(\frac{r}{R}\right)^2}$$

# General Geometry

$$dl^2 = r^2(d\theta^2 + \sin^2 \theta d\phi^2) + \frac{dr^2}{1 - \left(\frac{r}{R}\right)^2} \text{ (spherical)}$$
$$+ dr^2 \text{ (flat)}$$
$$+ \frac{dr^2}{1 + \left(\frac{r}{R}\right)^2} \text{ (hyperbolic)}$$

The Robertson-Walker line element:  $r \rightarrow \frac{r}{R}$

$$ds^2 = -c^2 dt^2 + R(t)^2 \left[ r^2(d\theta^2 + \sin^2 \theta d\phi^2) + \frac{dr^2}{1 - kr^2} \right]$$

with  $k = -1, 0, +1$  accordingly to geometry.

# General Geometry

The Robertson-Walker line element:  $r = r R_0$

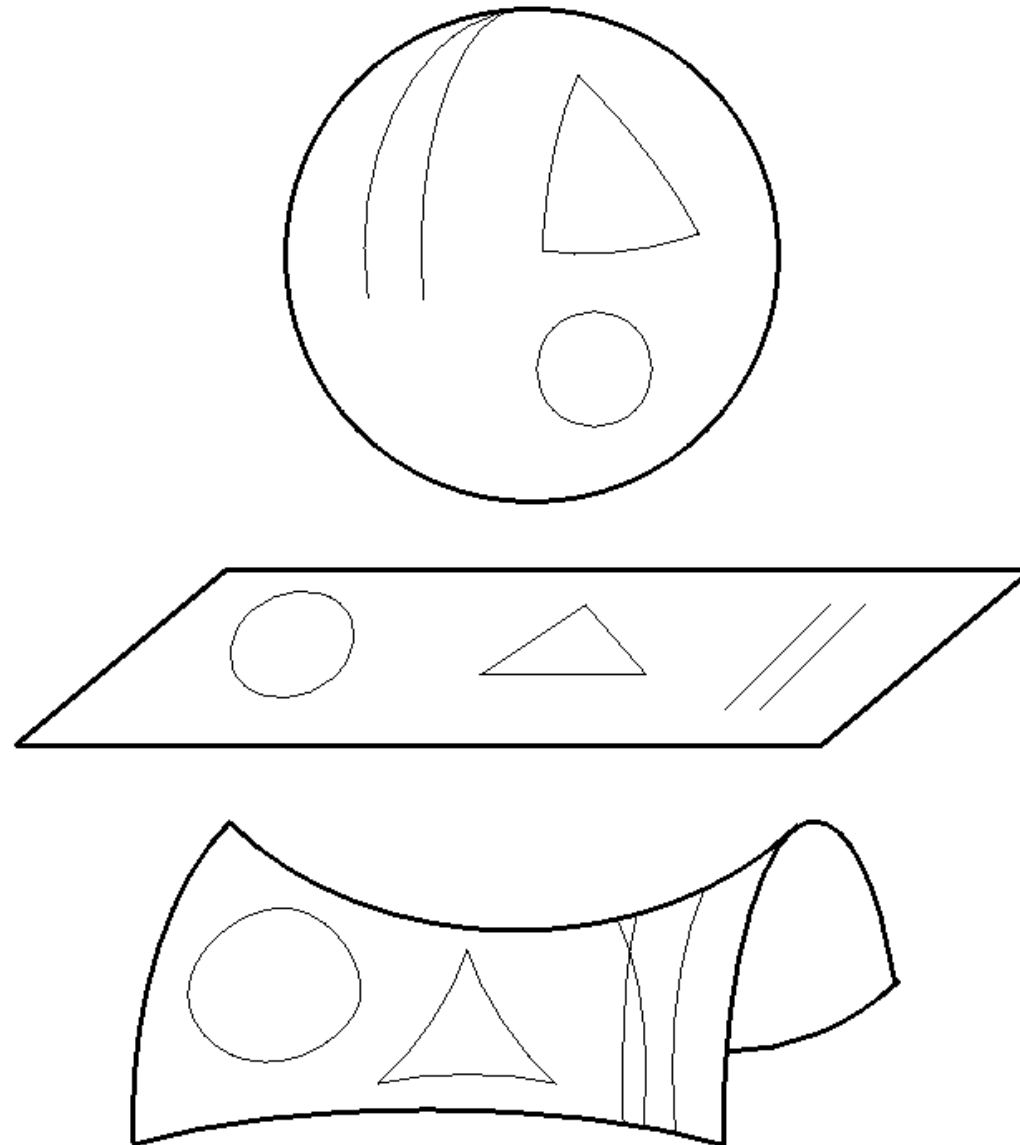
$$ds^2 = -c^2 dt^2 + a(t)^2 \left[ r^2 (d\theta^2 + \sin^2 \theta d\phi^2) + \frac{dr^2}{1 - Kr^2} \right]$$

with  $K = \frac{k}{R_0}$  and  $a(t_0) = 1$ .



# General Geometry

Three possible geometries:



# Topology

The local geometry of space (i.e. the value of  $k$ ) does not prejudice of the global shape of space i.e. its topology.

-> The universe could be finite even with  $k = 0, -1$ .